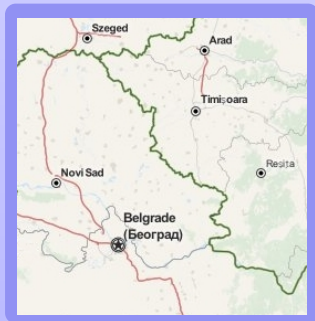


West University of Timișoara
Faculty of Physics

Aspects of quantum modes on de Sitter spacetime

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Faculty of Physics

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- thesis: "Contributions to the quantum field theory on de Sitter spacetime"
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Outline of the Short-Talk

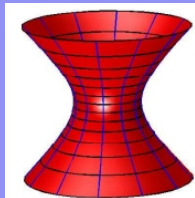
- 1 de Sitter spacetime
 - The de Sitter background
 - Charts on dS spacetime
- 2 Quantum Modes
 - de Sitter spacetime and modes
 - Conserved operators
 - Defining the quantum modes
 - Modes on the euclidean chart
 - Spherical energy basis modes
- 3 Concluding remarks

The de Sitter manifold

- can be embedded in a 5D Minkowski space

Constraint:

$$\eta_{AB}Z^AZ^B = -\frac{1}{\omega^2}$$



embedding

manifold	M^5	\longrightarrow	dS
coords.	$\{Z^A\}$		$\{x^\mu\}$
metric	η^{AB}	$Z^A = Z^A(x^\mu)$	$g^{\mu\nu}$

Metric tensor and Killing vectors

Induced metric on dS

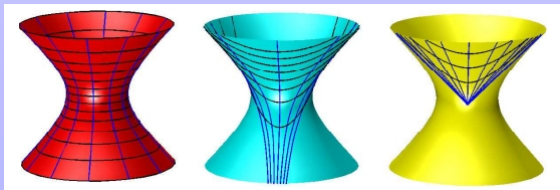
$$g_{\mu\nu} = \eta_{AB} \frac{\partial Z_A}{\partial x^\mu} \frac{\partial Z_B}{\partial x^\nu}$$

- inherits its isometry group from the gauge group of \mathbb{M}^5 : $SO(1,4)$
- is a maximally symmetric spacetime (has 10 Killing vectors)

$$k_{AB}^\mu = g^{\mu\nu} Z_A \overset{\leftrightarrow}{\partial}_\nu Z_B$$

FLRW charts

- dS manifold is the only one that admits all 3 types of FLRW charts
- important for cosmology: exhibit isotropy (rotational symmetry is manifest) and homogeneity
- ∂_t is not a Killing vector



$k = +1$
hyperspherical

$k = 0$
spatially flat

$k = -1$
hyperbolic

¹Moschella, Progr.Math.Phys.47:120 (2006)

static chart $\{t_s, r_s, \theta, \phi\}$

$$ds^2 = (1 - \omega^2 r_s^2) dt_s^2 - \frac{dr_s^2}{1 - \omega^2 r_s^2} - r_s^2 d\Omega_2^2$$

- ∂_{t_s} is a Killing vector

dS- Painlevé chart $\{t, r_s, \theta, \phi\}$

$$ds^2 = (1 - \omega^2 r_s^2) dt^2 + 2\omega r_s dr_s dt - dr_s^2 - r_s^2 d\Omega_2^2$$

- hybrid between static chart and FLRW chart
- time slices are euclidean spaces
- ∂_t is a Killing vector

'Natural' charts

- time and space on equal footing $Z^\mu = \frac{x^\mu}{f(x)}$
- basis for a so-called 'de Sitter-invariant special relativity':
- Beltrami chart ³: $f(x) = \sqrt{1 - \omega^2(t^2 - \vec{x}^2)}$ - the "inertial coordinates" of dS
- stereographic chart ⁴: $f(x) = 1 - \omega^2(t^2 - \vec{x}^2)/4$ - conformal to Minkowski spacetime
- NO symmetries are manifest

³Guo, Huang, Xu, Zhou, Mod.Phys.Lett.A19:1701 (2004)

⁴Aldrovandi, Almeida, Pereira, Class.Quantum Grav.24:1385 (2007)

Quantisation of fields on de Sitter spacetime

- gravity remains classical, only matter fields are quantized!
- dS background - arena of interactions for the quantum theory (i.e. fields do NOT interact with the background)

Minkowski spacetime \rightarrow deSitter Spacetime

- no. of symmetries remains the same! (10 Killing Vectors) - give rise to conserved operators
- parameter introduced: ω (related to the cosmological constant). In the limit $\omega \rightarrow 0$, Minkowski quantities should be obtained
- the complexity of equations and their solutions increases
- first step: determine the free fields on the dS manifold

The conserved operators

Scalar conserved operators:

$$X_{AB} = -ik_{AB}^{\mu} \partial_{\mu}$$

Hamiltonian operator:

$$H = \omega X_{04}$$

Angular momentum:

$$J_i = i\varepsilon_{ijk} X_{jk}$$

Momentum operator:

$$P_i = \omega(X_{i4} + X_{0i})$$

'Runge-Lenz-type'
operator:

$$R_i = X_{i4}$$

The 'correct' momentum operator on dS

Minkowski limit:

$$\lim_{\omega \rightarrow 0} P_i = \lim_{\omega \rightarrow 0} \omega R_i = P_i^{\text{M}} \equiv -i\partial_i$$

Commutation relations:

$$[P_i, P_j] = 0$$

$$[R_i, R_j] = i\epsilon_{ijk} J_k$$

- P_i are good momentum operators

Also:

$$[H, P_i] = i\omega P_i$$

- Energy and momentum can't be measured simultaneously on dS

- The field operator can be expanded as $\Phi(x) = \int da db dc f_{abc}(x)a(a,b,c) + f_{abc}^*(x)a^\dagger(a,b,c)$, and must satisfy:

Field equation- $s=0$ (Klein-Gordon equation)

$$\frac{1}{\sqrt{|g|}} \partial_\mu (\sqrt{|g|} g^{\mu\nu} \partial_\nu \Phi(x)) - m^2 \Phi(x) = 0$$

- AND

Eigenvalue equations for operators:

$$A\Phi(x) = a\Phi(x), \quad B\Phi(x) = b\Phi(x), \quad C\Phi(x) = c\Phi(x)$$

- The latter give the separating constants a physical meaning: quantities arising from measurements corresponding to a

C.S.C.O. (complete set of commuting operators)

$$\{E, A, B, C\}$$

Spatially flat FLRW (euclidean) chart

Metric

$$ds^2 = dt^2 - e^{2\omega t} d\vec{x} \cdot d\vec{x}$$

Conserved operators:

- spatial translations are manifest

$$P_i = -i\partial_i$$

- time translation accompanied by a spatial dilatation (in accordance with the concept that dS spacetime is expanding)

$$H = -i(\partial_t + x^i \partial_i)$$

- P_i do not commute with $H \Rightarrow$ two kinds of mode functions

De Sitter spacetime with $ds^2 = dt^2 - e^{2\omega t} d\vec{x} \cdot d\vec{x}$

-momentum-basis plane waves ⁵:

$$f_{\vec{p}}(t, \vec{x}) = \frac{1}{2} \sqrt{\frac{\pi}{\omega}} \frac{1}{(2\pi)^3} e^{-\frac{3\omega t}{2}} e^{\frac{i\pi\nu}{2}} H_{\nu}^{(1)}\left(\frac{p}{\omega} e^{-\omega t}\right) e^{i\vec{p} \cdot \vec{x}}$$

-energy-basis plane waves ⁶:

$$f_{E, \vec{n}}(t, \vec{x}) = \frac{1}{2} \sqrt{\frac{\omega}{2}} \frac{1}{(2\pi)^3} e^{-\frac{3\omega t}{2}} e^{\frac{i\pi\nu}{2}} \int_0^\infty \sqrt{s} H_{\nu}^{(1)}(s e^{-\omega t}) e^{i\omega s \vec{n} \cdot \vec{x} - i \frac{E}{\omega} \ln s}$$

⁵Nachtmann, Commun.math.Phys.6:1 (1967)

⁶Cotăescu, Crucean, Pop, Int.J.Mod.Phys.A23:2563 (2008)

The 'Schrödinger Picture' formalism⁷

Apply a (possibly non-unitary) operator $U(x)$

$$\Phi(x) \rightarrow \Phi_S(x) = U(x)\Phi(x)$$

$$O \rightarrow O_S = U(x)OU(x)^{-1}$$

where

$$U(x) = e^{-\omega t(x^i \partial_i)}$$

such that

$$U(x)F(x^i)U(x)^{-1} = F(e^{-\omega t}x^i)$$

$$U(x)F(\partial_i)U(x)^{-1} = F(e^{\omega t}\partial_i)$$

⁷Cotăescu, arXiv:0708.0734 (2007)

Deriving the equation

Klein-Gordon eq. in $\{t, x, y, z\}$ chart

$$(\partial_t^2 + 3\omega\partial_t - e^{-2\omega t}\Delta_{x,y,z} + m^2)\Phi(t, \vec{x}) = 0$$

Natural Picture \rightarrow Schrödinger Picture:

$$\partial_t \rightarrow \partial_t + \omega x^i \partial_i$$

$$\partial_i \rightarrow e^{\omega t} \partial_i$$

$$\Delta \rightarrow e^{2\omega t} \Delta$$

$$\Phi(t, \vec{x}) \rightarrow \Phi_S(t, \vec{x})$$

Klein-Gordon eq. in $\{t, x, y, z\}$ chart - Schrödinger Picture

$$((\partial_t + \omega x^i \partial_i)^2 + 3\omega(\partial_t + \omega x^i \partial_i) - \Delta_{x,y,z} + m^2)\Phi_S(t, \vec{x}) = 0$$

$$\{t, \vec{x}\} \rightarrow \{t, r, \theta, \phi\}:$$

$$x^i \partial_i = r \partial_r$$

$$\Delta_{x,y,z} = \Delta_{r,\theta,\phi} = \partial_r^2 + \frac{2}{r} \partial_r + \frac{\Delta_{\theta,\phi}}{r^2}$$

Klein-Gordon eq. in $\{t, r, \theta, \phi\}$ chart - Schrödinger Picture

$$\left((\partial_t + \omega r \partial_r)^2 + 3\omega(\partial_t + \omega r \partial_r) - \partial_r^2 - \frac{2}{r} \partial_r - \frac{\Delta_{\theta,\phi}}{r^2} + m^2 \right) \Phi_S(t, r, \theta, \phi) = 0$$

Solution (back in Natural Picture)

$$f_{E,l,m_l}(t, r, \theta, \phi) = N e^{-iEt} (\omega r e^{\omega t})^l {}_2F_1(\alpha, \beta; l + 3/2; \omega^2 r^2 e^{2\omega t}) Y_{l,m_l}(\theta, \phi)$$

Solution- in integral representation (via Hankel transform)⁸

$$f_{E,l,m_l}(t,r,\theta,\phi) = N \times 2^{-i\epsilon} \frac{i\pi}{2} e^{\frac{i\pi\nu}{2}} \frac{\Gamma(l+3/2)}{\Gamma(\alpha)\Gamma(\beta)} Y_{l,m_l}(\theta,\phi) e^{-\frac{3\omega t}{2}} \frac{1}{\sqrt{\omega r}} \times \\ \times \int_0^\infty s^{-i\epsilon} H_\nu^{(1)}(se^{-\omega t}) J_{l+1/2}(\omega rs) ds$$

Scalar Product (in spherical coordinates)

$$\langle f_{E,l,m_l}, f_{E',l',m'_l} \rangle = i \int_0^\infty d^3r r^2 \int d\Omega e^{3\omega t} f_{E,l,m_l}^*(t,r,\theta,\phi) \overleftrightarrow{\partial}_t f_{E',l',m'_l}(t,r,\theta,\phi)$$

$$N = \sqrt{\frac{\omega}{2}} \frac{1}{\pi} \frac{\Gamma(\alpha)\Gamma(\beta)}{\Gamma(l+3/2)}$$

Concluding remarks

- the limiting case $\omega \rightarrow 0$ can't be evaluated for the $k=0$ FLRW chart quantum modes, but that's OK: modes are not observable quantities
- on a chart there can be more than one useful mode-expansion of the field operator
- while the energy-basis modes are different from the momentum-basis ones, there is still no Bogolyubov mixing (vacuum is stable under transf. from one set to another)
- the most useful charts for computing quantum modes are the ones where symmetries are manifest
- spatially flat FLRW chart- translational symmetries are manifest $\Phi(x) \sim e^{i\vec{p}\cdot\vec{x}}$ - use of free modes in a QFT with perturbative Feynman-Dyson formalism⁹

⁹Crucean, Mod.Phys.Lett.A22:2573 (2007)

Cotăescu, Crucean, Int.J.Mod.Phys.A23:1351 (2008)

Thank you for
your attention!